

## Re: Swift grb satellite

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*Source:* <http://sci.tech-archive.net/Archive/sci.astro/2005-04/msg00163.html>

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- *From:* [jaymoseley@xxxxxxxxxxxx](mailto:jaymoseley@xxxxxxxxxxxx) (sean)
  - *Date:* 27 Apr 2005 01:49:50 -0700
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Craig Markwardt <[craigmnet@xxxxxxxxxxxxxxxxxxxxxxxxxxxxxxxx](mailto:craigmnet@xxxxxxxxxxxxxxxxxxxxxxxxxxxxxxxx)> wrote in message news:<[ony8b8v5hz.fsf@xxxxxxxxxxxxxxxxxxxxxxxx](mailto:ony8b8v5hz.fsf@xxxxxxxxxxxxxxxxxxxxxxxx)>...

From: Craig Markwardt ([craigmnet@xxxxxxxxxxxxxxxxxxxxxxxxxxxxxxxx](mailto:craigmnet@xxxxxxxxxxxxxxxxxxxxxxxxxxxxxxxx))  
Subject: Re: Swift grb satellite

>> I required them to be released.

>Irrelevant. Your requirements are not part of the Swift mission.

I also note that you ignored the fact that light curves and spectra  
\*are\* released via GCN notices, and in some already-published papers.

What gcn notices? I have yet to see one spectra released to the public  
And what papers? It doesnt mention in the SWIFT website  
that the spectra will be available in only certain papers. Let  
alone which papers these will be available in.

>I note that you deleted \*yet again\* an example which shows the  
>contradictory nature of your two claims. You could have researched  
>the definition of "proportional" but did not.

Please supply the `deleted` example that supposedly shows which  
of the two does \*not\* mean that the delay increases proportional  
to wavelength. In fact it is you fails to answer the above question  
I pose. Which of the two phrasings I use Does NOT mean that  
the delay increases proportional to wavelength?

>> The Lyman break is a sharp break whereas there is a

>This claim is not true, therefore the conclusions you draw based on it  
>are irrelevant.

Are you suggesting that the Lyman forest is seen on the long wavelength  
side of 1215? In case you forgot it starts at ~1215.

And you didnt answer my question.. The Lyman break is sharp at  
about 1215 whereas the observed drop off is gradual. There wouldnt  
be a gradual drop off with decreasing wavelength if it were due to  
the Lyman alpha forest. So please tell me why does this gradual drop

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off not match a sharp Lyman break?

><http://mathworld.wolfram.com/Proportional.html>

> If 'a' is (directly) proportional to 'b', then 'a/b' is  
> constant. The relationship is written , which implies 'a = c b' for  
> some [proportionality] constant 'c'.

>The proportionality constant between wavelength (\*) and time is the  
>adjustable parameter of your model, whether you realize it or not.

>(\*) – or difference in wavelength, for your other contradictory claim.

So if `a` is the peak observed at wavelength `a` , lets say 20nm,  
and `b` is the observed time of `a` peaking at 1 second,.. then in  
another wavelength `a` of 200nm, the delay `b` would be in my model  
10 seconds.

Thats  $20/1 = 20$  and  $200/10 = 20$ .

Whats your problem? They both are the same result of 20,  
or in other words constant. Your whole argument is proven wrong  
by my calculation here.

>I didn't ask for a  $\chi^2$  fit to test your model. What I did suggest  
>is that some kind of formal statistical test which addresses  
>measurement uncertainties, upper limits, and derives parameter  
>confidence regions, is required. I note your diversion to the  
>"smoothing" topic.

And I ask you to be more specific. What sort of formal tests  
(upper limits etc) do you require me to do? Specify please and  
I will deliver. I cant read your mind.

>"Systematic bias" has a well-defined meaning, and addressing it is  
>.crucial to most scientific research.

>That would be your problem, not mine. There are obvious selection  
>effects which would lead to biases. For example, optical/IR  
>observations can usually be done the night of the GRB, but as a  
>practical matter, sensitive radio observations often take longer to  
>schedule. This would naturally lead to a gamma-ray – then optical/IR  
>– then radio progression of detections. But of course this only  
>reflects the order of \*observations\*, not the presence of emission,  
>nor a "peak" in the emission.

OK this is a bit more specific regarding systematic biases. Thanks.  
But it doesnt seem to say anything that impinges on my model. For  
instance the systematic biases cited above would only reflect  
on when and how long the observations take place, and how frequently.  
Yet this has no bearing on my model in that if there are less  
scheduled observations in radio you say that means that  
the ones that are available are not valid. Are you are suggesting that

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if there are less radio observations for one grb than the ones that have been made are less accurate?

The only relevance would be that you are saying that if there are less observations in one wavelength the less clear the decay rate is. That's true but it doesn't mean that if I predict that the decay rate will follow a precise lightcurve that these fewer observations won't follow that lightcurve. Because in fact so far, whether it's 100 or 2 observations taken under widely different circumstances those 2 or 100 still can be shown to fall within the parameters set out by my predictions. And you have yet to supply 1 grb out of all in history that doesn't.

>Yes, and for a century, it has been known that electromagnetic >radiation travels at a constant speed at all wavelengths. Thus your >"model" is utterly erroneous.

And this is the untruth I mean to expose. Because the whole point of this theory is to show that in grbs it is the first provable case where light can be not only shown to travel at different speeds relative to the observer but can ONLY be explained by  $c$  being variable. Why else would I dwell on this if it wasn't that I felt my model was the first indisputable evidence of a variable  $c$  where modern physics could not offer an equal explanation

>Of course, if as you claim, that GRBs are simply upshifted starlight, >then the spectra in other wavelength bands should be stellar spectra >— BUT THEY ARE NOT. In fact, gamma-ray burst spectra have well known >bent power law spectra (stars do not). X-ray spectra of afterglows >are power laws (for stars they are not). The optical/IR spectra of >GRB afterglows do not match optical/IR spectra of stars. Stellar >spectra typically have systems of absorption and emission lines >superimposed on a thermal continuum — GRB spectra at all wavelengths >are not like this. Thus your model is multiply erroneous.

You don't make sense here. For instance your first point.

1. "the spectra in other wavelength bands should be stellar spectra" what does this mean? What spectra? GRB spectra or local stars spectra?

2. You misunderstand my model. How can you criticize it if you don't understand it? For instance you say all the above grb observations are not the same as star spectra! Of course not. Am I saying that grbs are local stars! No. I say that grbs are star spectra from great cosmological distances  $> 15 - 20$  billion years at least that are being redshifted \*as we watch.\*

So that one minute this distant stars spectra is observable let's say between 10–70nm and then 10 seconds later it's redshifted so much so that it's stretched to 100–700nm etc. Are local stars spectra from  $< 10$  billion light years being redshifted from gamma to optical in minutes as we watch??? No of course not. That's why grb's look different., that's why I call them optical illusions.

So you saying above that grb are different from local spectra only

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reaffirms how correct my model is, because that's what I have been arguing all along. In fact it's actually good news in a way that you above are actually beginning to understand the difference between light from a nearby star and light from a very distant rapidly redshifting star (grb) in a non BB non GR universe.

You actually admit now that grbs are not like star sources. Obviously you have studied these embargoed spectra and realize I'm right.

>>I note that you did not actually read astro-ph/0503508 or  
>>astro-ph/0503521, which both show \*observed\* prompt optical  
>>emission for GRBs 990123 and 041219a.

>

> Your two papers speculate that prompt optical emission may  
> have occurred but the fact is that there is no observation  
> that was made in optical at the time of trigger in gamma.

>This is an erroneous claim. In fact, for GRB 041219a, optical  
>observations were taken during the peak of the gamma-ray burst, and  
>optical emission was detected (see citations given). Thus, "zero"  
>delay between gamma-ray and optical wavelengths is possible.

Yes but was optical seen to peak at the same time as the peak of gamma?

How many optical observations were made during the gamma observed part of the burst? Enough to confirm that an observation in one filter band was seen to decay in other observations in the same filter band while gamma was still observable?

I don't think so and if you think so supply me with those gcn because I have already supplied gcn info that shows this is not an observed fact (that an optical band was seen to increase and peak and then decrease at the same time as gamma.) No such data is available and you cannot supply any.

>Ironic, since in the case of 041219a, the optical emission peaked at  
>the time of the gamma-ray peak. It looks like the preclusion of your  
>model is incorrect.

I'm not sure if we mean the same thing here or are talking across each other. If you mean prompt optical emission is any optical OT seen while gamma is still being observed then I don't dispute that nor does my theory. What I'm saying is not possible is that in 041219a the optical could not be observed to peak at the same time as gamma but always later. So it can in an unusually long burst like 041219a, be seen to overlap but never will optical peak at the same time as gamma. And the observations support my claim.

K was seen to peak AFTER gamma was faded. K is considered optical I assume so therefore it's not true to say that optical peaked at the same time as gamma in 041219a. K contradicts this.

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>> exposures that the decay rate is a straight  
>> line between two adjacent datapoints when in many cases the  
>> decay rate between any three exposures is variable.

>Pretense is not needed. Your claim is that the optical light curve  
>should be a stretched version of the gamma-ray light curve. In fact,  
>most of the light curves shown in the Berger paper are very well  
>sampled -- sampled enough to show a stretched version of the gamma-ray  
>light curve -- and look nothing like the gamma-ray light curve.

I dont agree . By well sampled I mean every 10–100 seconds not  
1 a day.

And exposure times should be 1–10 seconds not 500 as they seem to  
usually be.

What happens if in optical it is known already and I predict that  
the peak usually occurs over 1–10 hours? Most observations in  
any `1 band are made every day or every second day! Youll never  
get any detail.,THats why I Have made  
such a fuss about SWIFT. Because I assumed it would give  
observations in the first few hours in 3 or so different wavelengths  
and with each band having at least 5–10 seperate 16mag observations  
.. So far it seems unfortunately that SWIFT hasnt done this.  
You could though if next bright grb you took rotating exposures of  
10 seconds in each band followed by 10 seconds in the next like  
so... 10 s U, 10 s B, 10 s V, 10 s, U 10 s, B 10 s, etc  
so that in the firts minute you have 2 U, 2 B and 2 V etc etc .  
Then co add them all.Then coadd the first 8 V, etc, then co add  
the first 4 V etc then the first 2 V etc.  
Rather than 200 s U or 500 s V . Its too slow! The movies over  
before you even started watching it.

>Your analogy is irrelevant. In fact, the exposure times are short  
>compared to the time span of the overall light curves. This is  
>especially true on a logarithmic time axis. The optical light curves  
>are well enough sampled to detect a stretched version of the gamma-ray  
>light curve. With the existing sampling, there is really no way for  
>an optical light curve which decays by many orders of magnitude to be  
>mistaken for a gamma ray burst light curve.

I dont agree. Where have you an example of a single band, lets say  
600nm band where you have 20 seperate exposures of 30 seconds each  
made starting 5 minutes post burst with each 20 second exposure  
starting immediately after the last? For that matter where have you  
any example of the same frequency and exposure length at any time  
post burst in one band width?

In fact the only one I know of is one by the AAVSO observor Arno  
(ithink his name is or arto)and he did exactly this in two bands  
for 303329 and saw a peak clearly in one wavelength  
and it appeared delayed by a few minutes in a slightly longer  
wavelength observed band.

This observtion has still to be explained by anyone except me.

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>Again irrelevant. The whole point is for the IPN to make an  
>independent solution with the timing data only. Professional science  
>is not elementary school, and there is no need to manipulate the  
>results. The physics and geometry that underlie the IPN technique are  
>very simple.

OK it shouldn't be a problem for you to give me those 21  
grb numbers if you think I'm wrong. Let me analyse them and I'll  
show you how the success had larger overlaps than the failures  
on average.

>I note your continued lack of substantiation on your manipulation  
>claim. You could have analyzed the Laros data to substantiate your  
>claims of manipulation, but did not.

>> And as far as I am aware  
>> no OT was ever made of a grb till after these dates so how can  
>> this paper supply time of arrival locations that can be verified  
>> separately by Optical or radio observations?

>Irrelevant. Both Laros et al. papers show that IPN solutions are  
>consistent with GRO error boxes, and show that IPN solutions are  
>consistent with each other (in cases where multiple satellites are  
>involved). Your "model" would predict neither of those facts.  
>The locations of the IPN spacecraft are provided in each Laros et  
>al. paper, and they are given in the same coordinate system as the GRB  
>localizations. Thus your preconditions have already been met. The  
>rest is your problem.

I've looked at a couple grbs data from one of your cited papers and it  
seems there is enough to recalculate annulus. Although it would take  
a bit of study or help and probably a powerful computer (seeing as the  
98 paper points out the method was initially devised by a Los Alamos team  
of physicists presumably using more than an ipod) to correctly  
calculate an annulus. I'd like to point out that the 98 paper  
actually states that part of the process involves correcting IPN  
localizations by EYE(!) to make them work.

(Laros quote)

...."Furthermore, all of the quantitatively obtained results are checked  
for compatibility with eye estimates..."

Not what I would call a reassuring statement considering I have  
always claimed they adjusted the annulus to fit other verifiable  
localizations

Anyways separate from a detailed annulus calculation I found I can  
actually do a rough check by working out which is the closest satellite  
and the farthest of the three by comparing the 3 RA Dec positions of the  
3 satellites and then doing a rough calculation where Ulysses and pvo  
are similar to see if Ulysses is closer to the IPN localization. A  
couple I tried did roughly affirm the annulus calculations ~.

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The problem is, I cant verify if the IPN position is correct without a small alternative non IPN error box solution or preferably a known OT location with exact coordinates.

The two papers are pre 96 data only when I believe the first OT was discovered for any GRB. So for me the papers are useless without seperate non IPN verifiable localizations to check against.

What I need is as a Laros paper that has satellite ephemeris for GRB`s where there are also verified OT localizations or at least small HETE/ibis type error boxes.

I havent downloaded the 97 paper yet but I imagine its similar sort of data to 98 and as far as I can see the 98 paper offers NO seperate error boxes even though it hints that Batse does give them . Is it because Batse error box/circle was very large (like 1/2 the sky?)

Anyways as I have mentioned I now can double check whether an accurate non IPN localization, preferably OT coordinates, is consistent with time of arrival ephemeris which is probably more than anyone on this newsgroup is able to do with the possible exception of you(are you able to calculate this without calculating an annulus? Can you calculate an annulus?)

So if I can find ephemeris data AND seperate non IPN localizations for at least 4 grbs,I will be able to settle this .

Sean

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• **References:**

- ◆ **Re: Swift grb satellite**  
◇ From: sean
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◇ From: Craig Markwardt
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