

Sounds of Star Death Near Middle C (Forwarded)

Source: <http://sci.tech-archive.net/Archive/sci.astro/2006-01/msg00251.html>

- *From:* Andrew Yee <ayee@xxxxxxxxxxxxxxxxxxxxxxxxxxxxx>
 - *Date:* Thu, 26 Jan 2006 13:36:50 -0500
-

University Communications
University of Arizona
Tucson, Arizona

Contact Information:
Adam Burrows, 520-621-1795

January 24, 2006

Sounds of Star Death Near Middle C
By Lori Stiles

Scientists have made the astonishing discovery that sound might drive supernovae explosions. Their computer simulations say that dying stars pulse at audible frequencies -- for instance, at about the F-note above middle C -- for a split second before they blow up.

Researchers in the 1960s began using computer models to test ideas about what, exactly, causes stars to explode. But mathematical simulations have so far failed to satisfactorily explain the inner workings of nature's most spectacular blasts.

Neutrinos -- subatomic particles widely thought to power supernovae explosions -- don't seem to be energetic enough to do the job, especially for more massive stars. More sophisticated models that include convective motion work a bit better, but not well enough.

Adam Burrows of The University of Arizona and colleagues at UA's Steward Observatory, Hebrew University, and Germany's Max Planck Institute (Potsdam) have developed computer models that simulate the full second or more of star death, from the dynamics of core collapse through supernova explosion. Their two-dimensional computer models allow for the fact that supernovae outbursts are not spherical, symmetrical events.

Sounds of Star Death Near Middle C (Forwarded)

A supernova is a massive star that has burned for 10 million to 20 million years and developed a hot, dense 'white dwarf' star about the size of Earth at its core. When the white dwarf reaches a critical mass (about 1.5 times the mass of the sun), it collapses and creates a spherical shock wave, all within less than half a second before the star would explode as a supernova.

However, in all the best recent simulations, the shock wave stalls. So theorists have focused their work on what might revive the shock wave into becoming a supernova explosion.

According to Burrow's new results, part of the problem is that other computer models don't run long enough. His team's detailed models involve a million steps, or about five times as many as typical models that calculate only the first few hundred milliseconds of supernovae events. Burrows team's simulations also characterize the natural motion of a supernova core, something that other detailed models do not.

"Our simulations show that the inner core starts to execute pulsations," Burrows said. "And they allow us to follow the development to explosion for a longer time than other models do. They show that after about 500 milliseconds, the inner core begins to vibrate wildly. And after 600, 700 or 800 milliseconds, this oscillation becomes so vigorous that it sends out sound waves. In these computer runs, it's these sound waves that actually cause the star to explode, not the neutrinos."

He added, "We were quite sure when we started seeing this phenomenon that we were seeing sound waves, but it was so unexpected that we kept rechecking and retesting our results."

The team has used their models to make billions of calculations on computer clusters in the UA astronomy department, at Berkeley's supercomputer center and elsewhere, checking their analysis for the past year. They are publishing the research in the *Astrophysical Journal*. Their research is funded by the National Science Foundation, the Department of Energy, and the Joint Institute for Nuclear Astrophysics.

The team got a clear picture of what likely happens by making movies from their simulations. Burrows has posted the movies on his Website at <http://zenith.as.arizona.edu/~burrows/briley>

Collapsing material falls lopsidedly onto the inner core and soon excites oscillations at specific frequencies in the simulations. Within hundreds of milliseconds, the inner core vibrations become so intense that they actually generate sound waves. Typical sound frequencies are about 200 to 400 hertz, in the audible range bracketing middle C.

"Sound also generates pressure, which pushes the exciting streams of infalling matter to the opposite side of the core, further driving the core oscillations in a runaway process," Burrows said. "The sound waves reinforce the shock wave (created by the collapsed star) until it finally

Sounds of Star Death Near Middle C (Forwarded)

explodes aspherically."

Burrows said that others who study supernova explosions in computer experiments will be skeptical of his team's results -- and should be.

"This is such a break from 40 years of traditional thinking that one should be cautious trumpeting it," he said. "Nevertheless, this is provocative and interesting. It would open up many new possibilities and perhaps solve a long-standing problem of what triggers supernovae explosions."

Movies and images on related Website:

* Computer movie -- supernova explosion
http://uanews.org/20060123_12170.html

IMAGES:

[Available at <http://zenith.as.arizona.edu/~burrows/briley/>]

PHOTO CAPTION: burrows.explosion1.tiff:

Isodensity contours for the core of an 11-solar-mass star in explosion. The remnant neutron star is the green "dot" in the center and the outer shells are just interior to the blast wave that has been launched. The scale is ~2000 kilometers on a side and the time is ~800 milliseconds after the bounce at nuclear densities of the collapsed core. The colors reflect the entropies on the shells, where entropy is a measure of heat content. (Taken from the work of Burrows et al. (2006), Ap.J. in press.)

PHOTO CAPTION: s25_ye_on_den_still.tga:

Isodensity contours, colored with entropy, of the supernova blast seen in the calculations of Burrows et al. (2006, Ap.J. in press). Similar to the figure above, this figure shows the accretion funnels still streaming onto the inner core (purple dot). It is these accretion streams that continue to excite the core oscillation that generates the acoustic power that eventually explodes the star in these simulations. Green is high entropy and brown is low entropy material.

MOVIE DESCRIPTION: s25.dual.mpeg:

Movies of the growth of the core oscillation and explosion of a 25-solar-mass star. The left panel is a movie of a color map that depicts the density of the matter (red: high density; blue: low density) in the inner 50 kilometers of the core starting just after bounce at nuclear densities. The arrows are velocity vectors that follow the flow. The oscillation of the core grows to become significant at late times. Note the presence at times of both dipole and quadrupole character. The right panel shows the same evolution, but with an entropy color map (red: high

Sounds of Star Death Near Middle C (Forwarded)

entropy; blue: low entropy) in the inner 200 kilometers. The white lines are isodensity contours. At late times the core is oscillating violently and the object is exploding.

.